International Space Weather Summer Camp 2021: How to turn measurements into an analytical model – at the example of VLF data

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Knowledge for Tomorrow

Aftermath



Model applications: AMELIE

- target: locate and quantify the fall effect
- strongest slope identifiable also with mean/filter
- extremes can make troubles as not unique
- smoothing may yield to delays
- \rightarrow prescribed function can be treated analytically

Possibilities:

- decompose into symmetric and anti-symmetric part
- compare influence of non-symmetries for different Tx-Rx links





Global Ionospheric Flare Detection System

- now cast detection of SIDs caused by solar flares using a ground-based VLF system
- amplitude and phase measurements of various VLF transmitters
- industrial PC + SDR + MiniWhip antenna











Effects of space weather:

- increasing radiation exposure
 → Oct. 1989: extreme solar storm would be deadly for astronauts in protective clothing
- energetic particles can destroy satellite electronics leading to outages in communication and navigation
 → Jan. 1994: interruption of canad. telecom satellite Anik-E2 over 5 months
- induction of voltage on terrestrial line systems
 → Mar. 1989: power outage in Quebec for 9 hours
- deceleration of space debris due to the heating of the thermosphere \rightarrow 1979: crash of the US space station Skylab







GIFDS

Ground-based system

- easier maintenance than satellites
- measures actual effect → protecting technological devices on Earth

VLF monitor

- widespread use from navigation
- cost-efficient to receive
- X-ray flares detectable

Ionosphere Monitoring and Prediction Center

- space weather observation and forecast
- TEC maps from satellite measurements
- integration of GIFDS data and warnings





PITHIA-NRF

Plasmasphere Ionosphere Thermosphere Integrated Research Environment and Access services: a Network of Research Facilities





Model applications: GIFDS

- target: real-time flare warning
 → smoothing delays inacceptable
- need several links to cover globe
 → different levels critical to fade
- size estimation for source and effect
 → day parabola interferes with slope meaning

Possibilities:

- normalisation by subtracting "expected" value

 → comparable levels of individual signals
 → averaging to ongoing compound signal
 → detect and down-weight bad links
- equalise flare slopes and peak heights





Other influences

- normal radiation level not constant
 - \rightarrow another adjustment over the solar cycle (when enough measurements)
- ionisation change during lightnings or eclipse
 → affected links can be down-weighted
 → known change might be added





0 obscuration in %

Maximum

25



What about the phase?

Observation:

• amplitude

→ has heavier small-scale oscillations (difficult to detect anomalous increases)

• phase

 \rightarrow results in a "cleaner" measurement

Problem:

• phase experiences a drift

➔ similar evaluation techniques, but harder to measure

5 Nov 2013: C8.0 Flare 11:51-12:01



